

Cross calibration of Suzaku/Chandra/XMM-Newton with PKS2155-304

M. Ishida (ISAS/JAXA)

M. Tsujimoto (PSU)

T. Kohmura (Kogakuin Univ.)

Observation log

Suzaku/XIS

No.	ObsID	Date	t_exp (ks)	Sensor	Window	Mode	Charge Inj.	Processing ver.
SX1	700012010	2005/12/02	64	0,1,2,3	1/8	3x3, 5x5	off	2.0.6.13
SX2	101006010	2006/05/02	39	0,1,2,3	1/4	3x3, 5x5	off	2.0.6.13
SX3	102020010	2007/04/22	12	0,1,3	1/4	3x3	on	2.0.6.13
SX4		2008/05/12		0,1,3	1/4			

CXO/HRC/LETGS

No.	ObsID	Date	t_exp (ks)	Processing ver.
CH1	N/A			
CH2	6923	2006/05/01	30	7.6.7.2
CH3	8379	2007/04/22	30	7.6.11

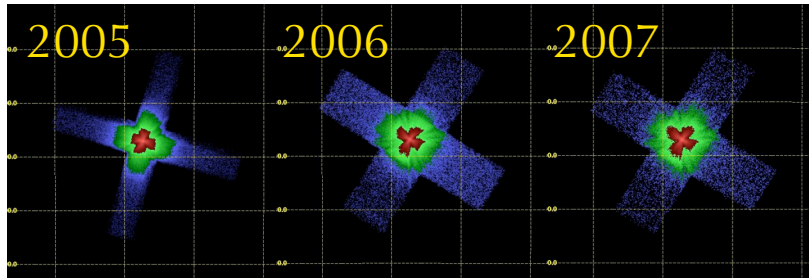
XMM-Newton/EPIC/MOS

No.	ObsID	Date	t_exp (ks)	Window	Filter	Position	Processing ver.
XM1	0158961301	2005/11/30	50	PrimePartialW2	Medium	Filter_C	CCF@2008/04/15
XM2	0158961401	2006/05/01	60	PrimePartialW2	Medium	Filter_C	CCF@2008/04/15
XM3	0411780201	2007/04/22	44	PrimePartialW2	Medium	Filter_C	CCF@2008/04/15

XMM-Newton/EPIC/pn

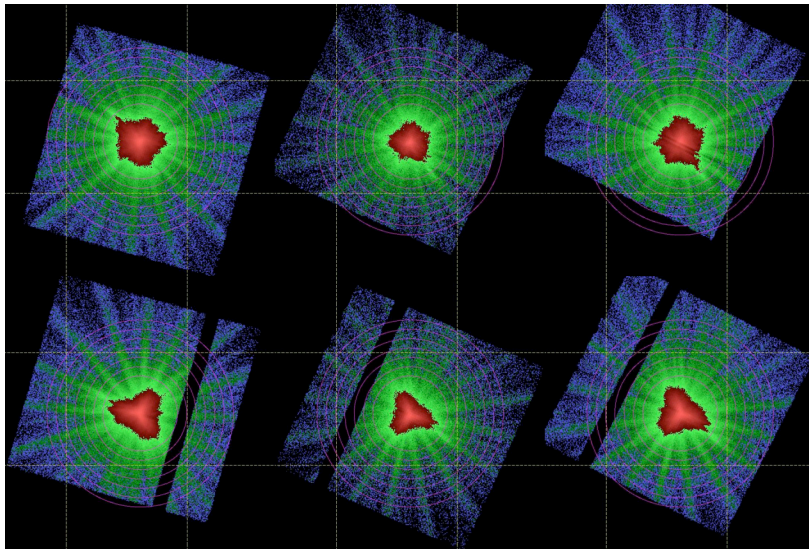
No.	ObsID	Date	t_exp (ks)	Window	Filter	Processing ver.
XP1	0158961301	2005/11/30	48	PrimeSmallWin	Medium	CCF@2008/04/15
XP2	0158961401	2006/05/01	52	PrimeSmallWin	Medium	CCF@2008/04/15
XP3	0411780201	2007/04/22	31	PrimeSmallWin	Medium	CCF@2008/04/15

Images



Suzaku XIS

2005: we do not use XIS1(BI) and XIS2 (FI) because significant part of their data suffer accidental light leak.



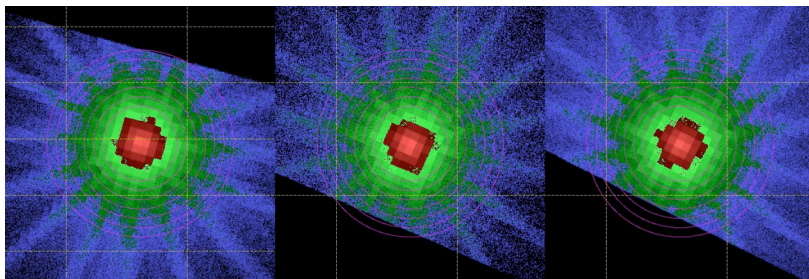
XMM/EPIC/MOS

Pileup at the image core.

2005: $r = 25''$ - $45''$

2006: $r = 25''$ - $40''$

2007: $r = 25''$ - $35''$



XMM/EPIC/pn

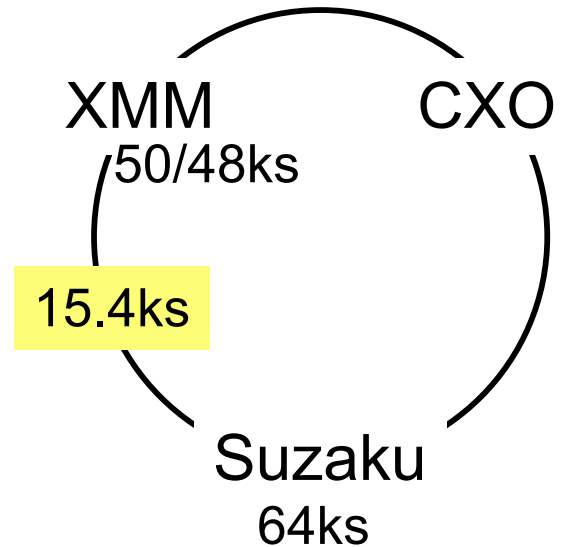
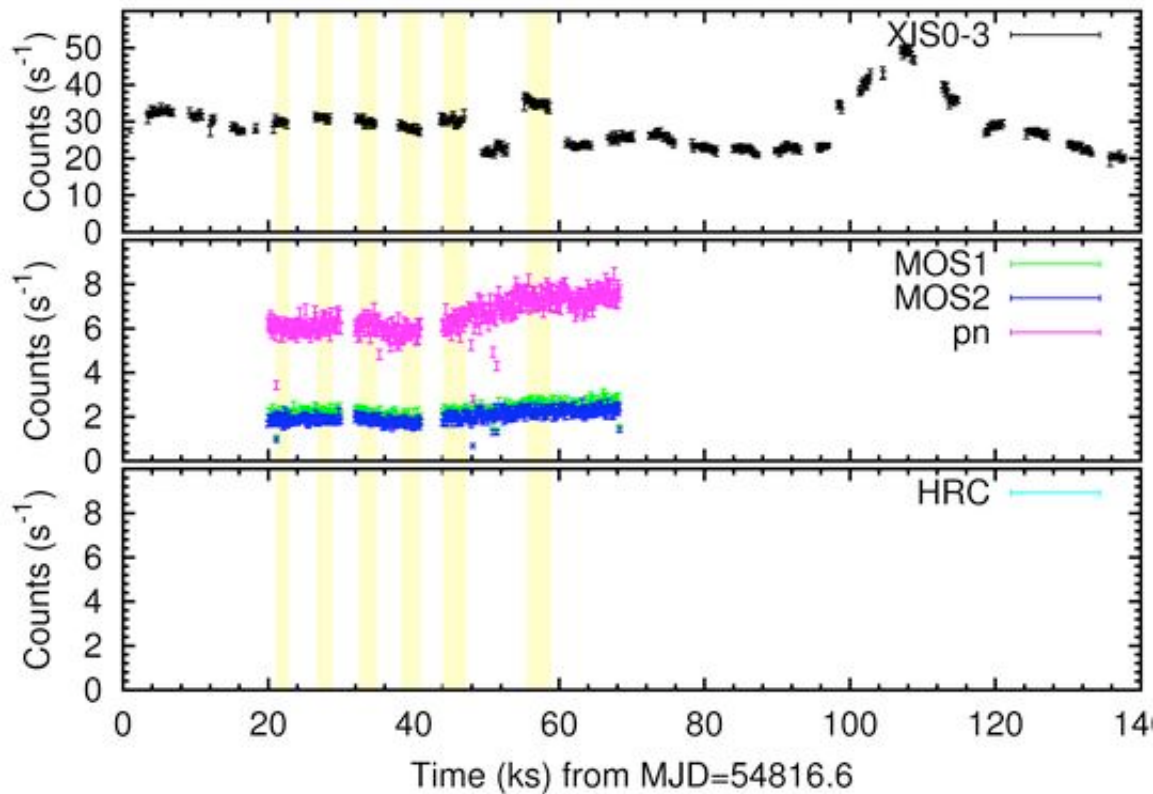
Pileup at the image core.

2005: $r = 25''$ - $45''$

2006: $r = 25''$ - $40''$

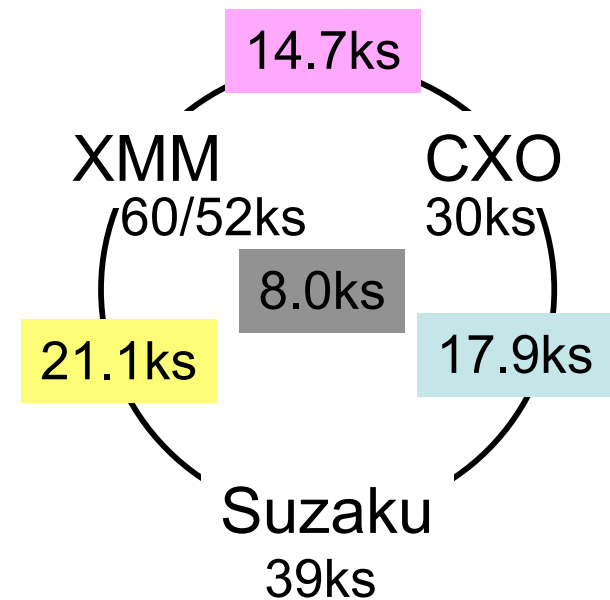
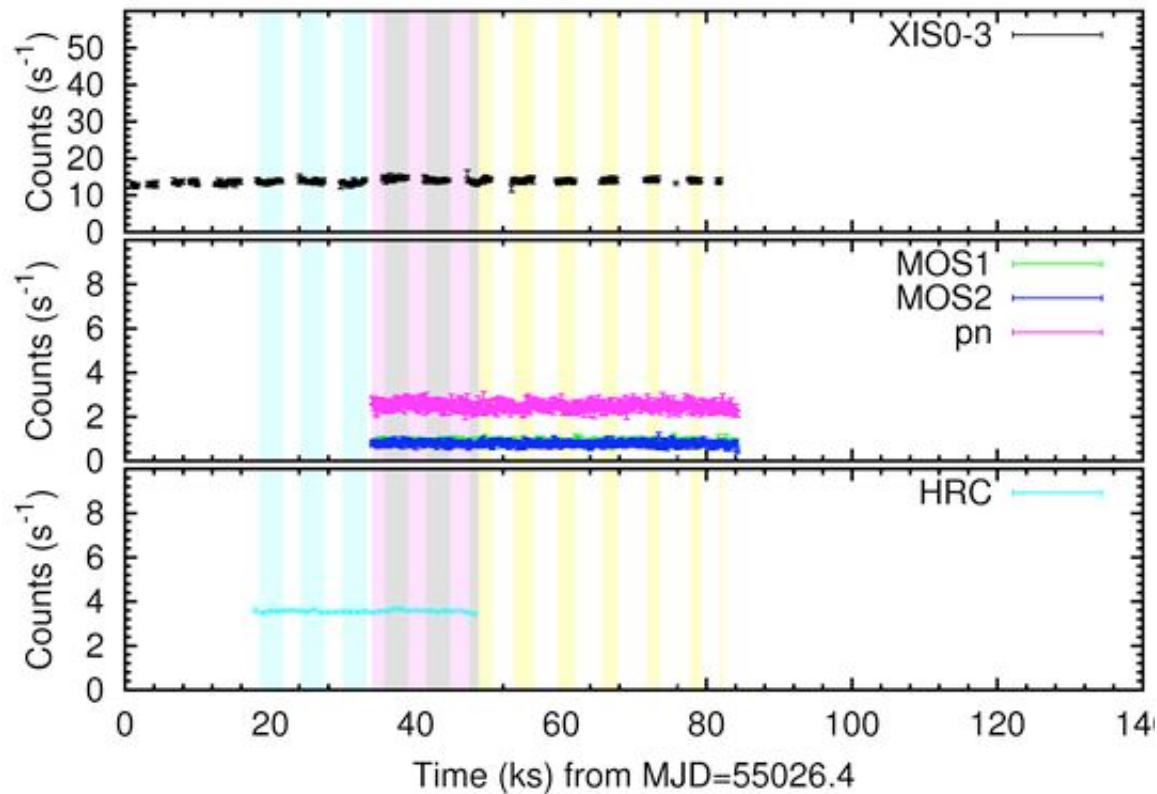
2007: $r = 25''$ - $35''$

Data overlap (2005)



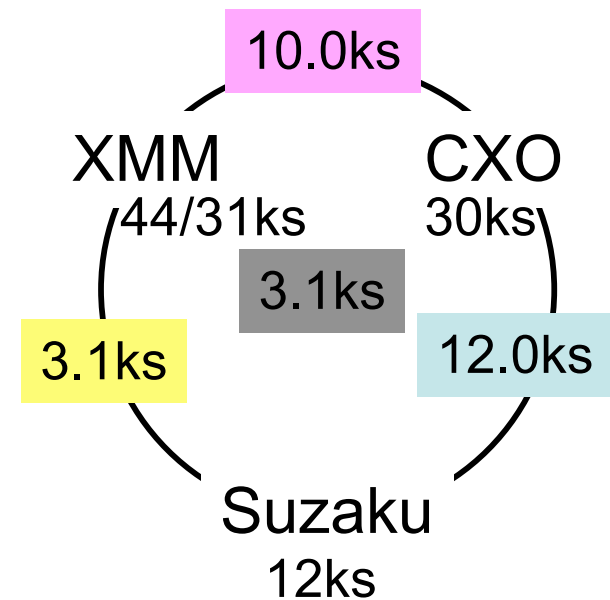
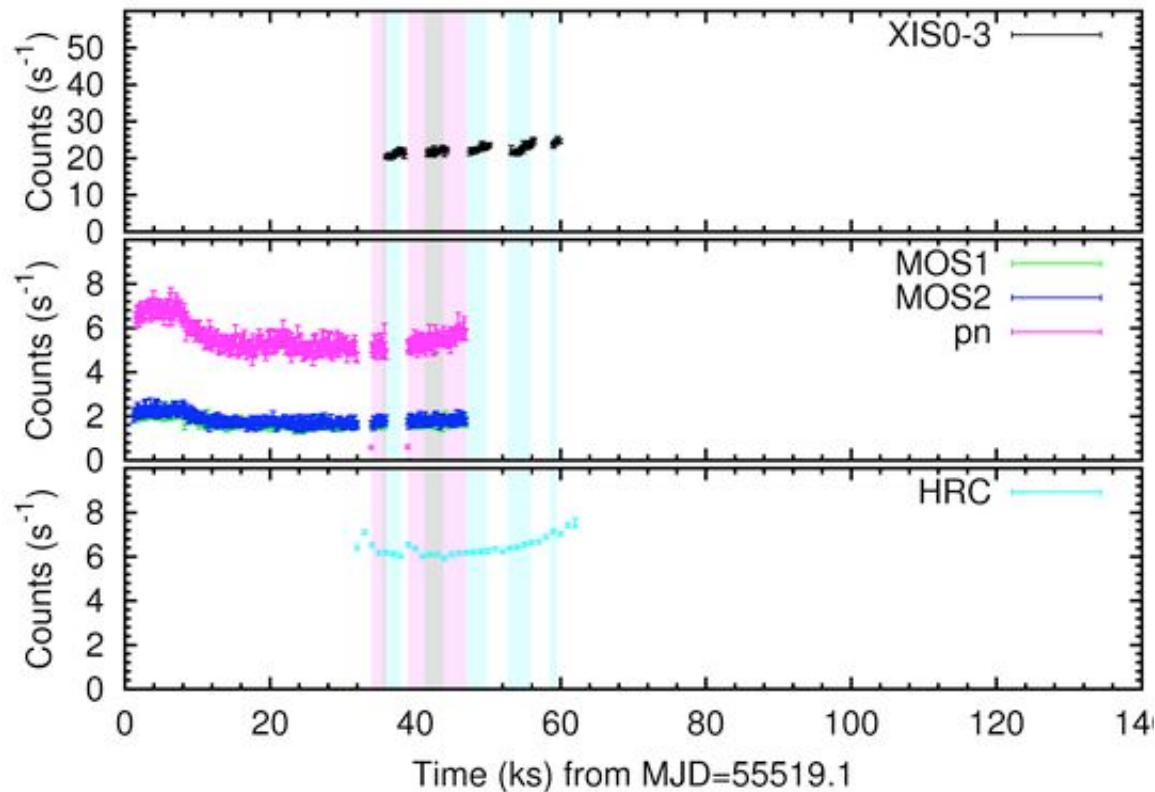
- The source was most intense among the three observations.
- No CXO coverage.
- Light leak of XIS1 and XIS2.

Data overlap (2006)



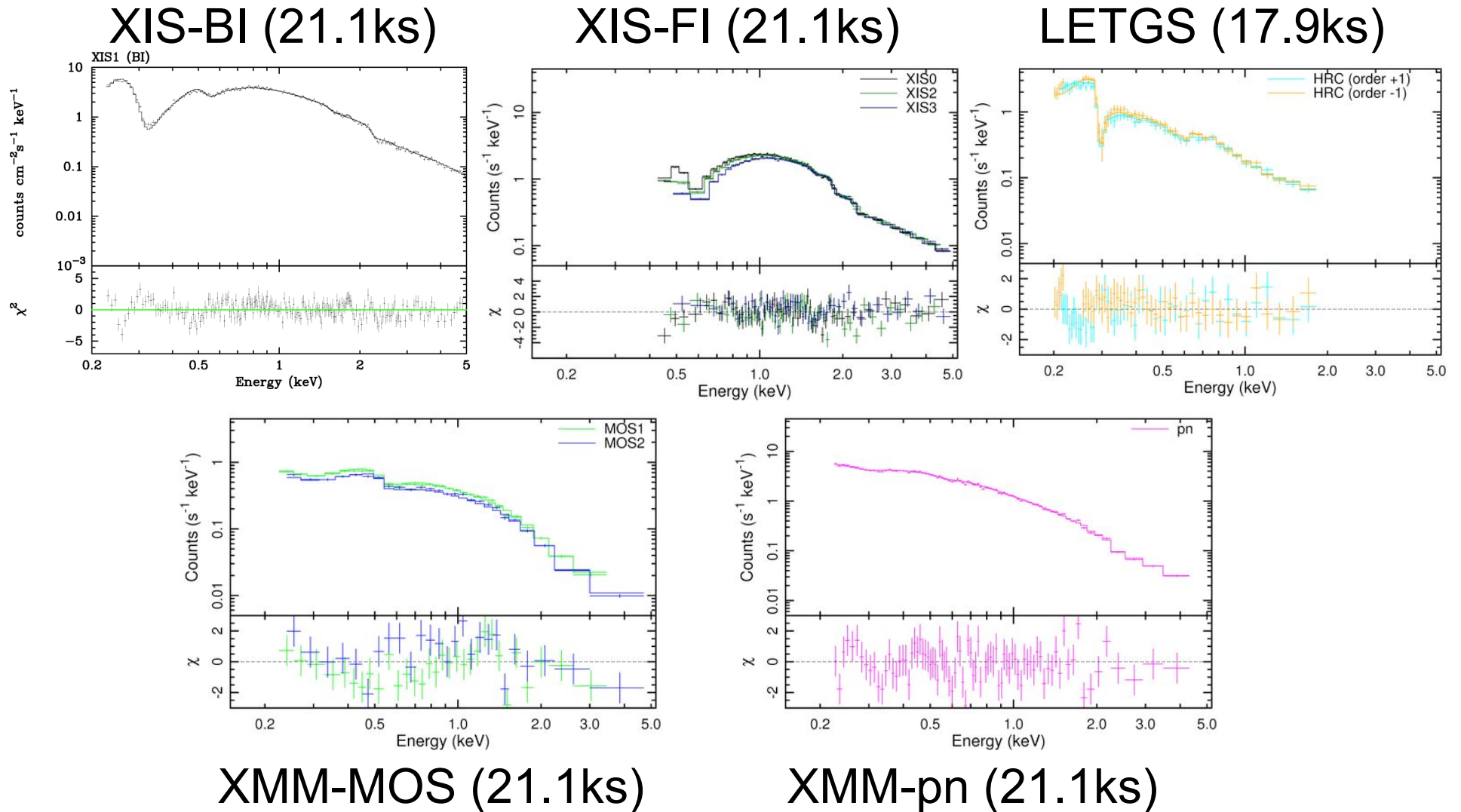
- The source was faintest among the three observations.
- Simultaneous coverage of any pair is moderate.
- Overlap of all instruments is limited.

Data overlap (2007)

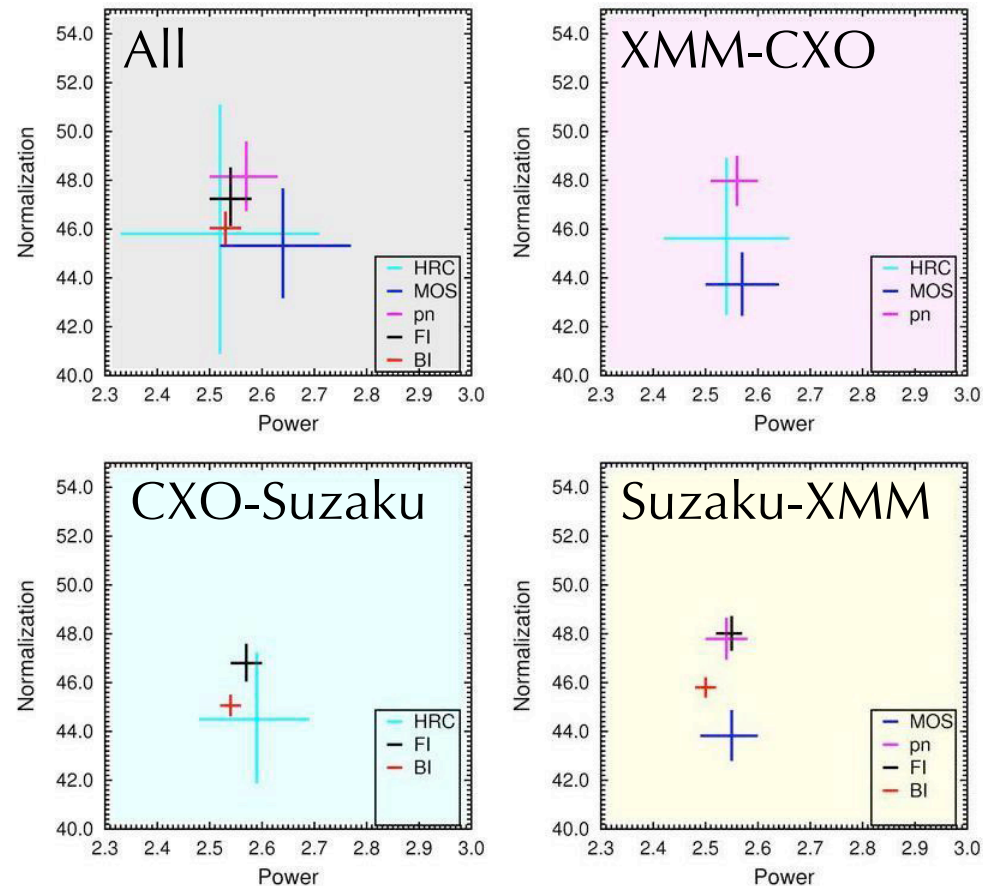


- The source intensity was intermediate among the three observations.
- Suzaku observations were too short.

Quality of Spectra (2006)

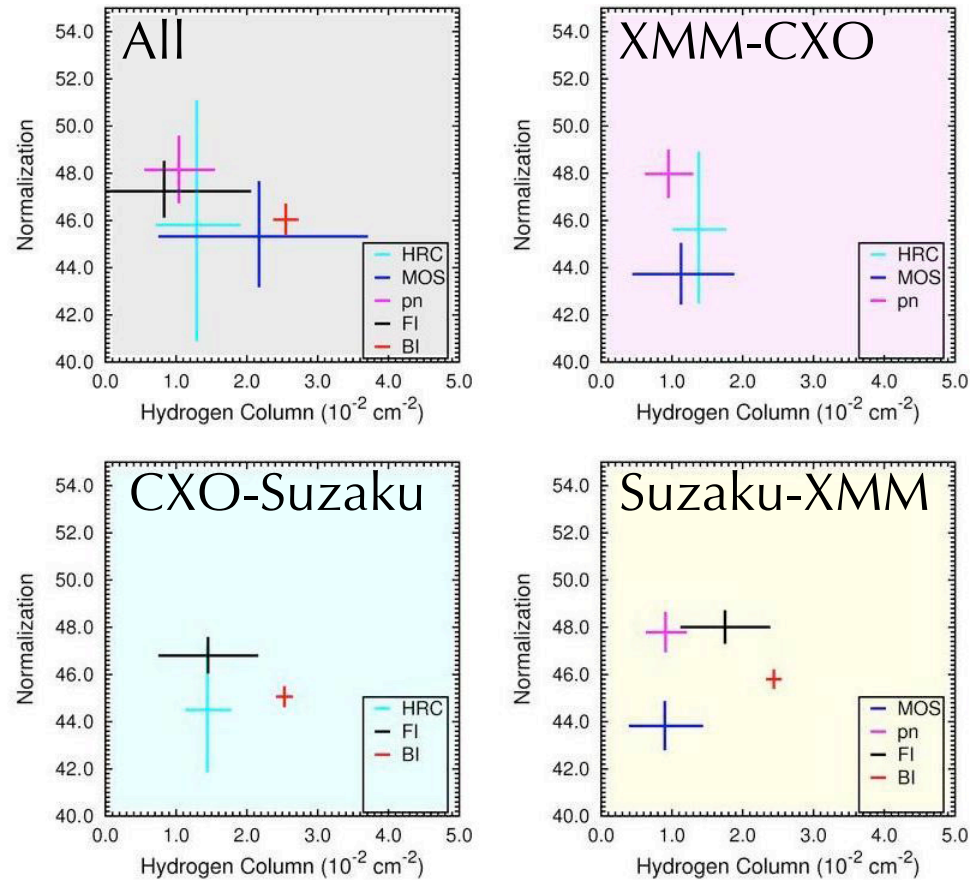


Comparison of photon index (2006)



- Photon index = 2.55 ± 0.03 for all pairs (ignoring errors).
- Normalization: flux in unit of $10^{-12} \text{erg cm}^{-2} \text{s}^{-1}$ in the 0.5-5keV band = $4.6 \pm 0.2 \times 10^{-11} \text{erg cm}^{-2} \text{s}^{-1}$ for all pairs (ignoring errors).

Comparison of N_H



- N_H to PKS2155-304 = $(1.5-1.7) \times 10^{20} \text{ cm}^{-2}$
- Uncertainty $\approx 1 \times 10^{20} \text{ cm}^{-2}$

Summary

- We carried out Suzaku/CXO/XMM coordinated observations of PKS2155-304 for cross calibration ~once per year.
- From the 2006 observations:
 - Flux in the 0.5-5keV agrees within $\pm 4\%$ ($4.6 \pm 0.2 \times 10^{-11}$ erg cm⁻² s⁻¹).
 - Photon index becomes within 2.55 ± 0.03
 - N_{H} distributes within $\approx 1 \times 10^{20}$ cm⁻² from the true value $(1.5-1.7) \times 10^{20}$ cm⁻².
- For better calibration (overlap time > 40ks),
 - CXO and XMM should expose >40ks, with maximum effort to overlap between the two.
 - Suzaku should cover both CXO and XMM observing windows (>20ks overlap with XMM).