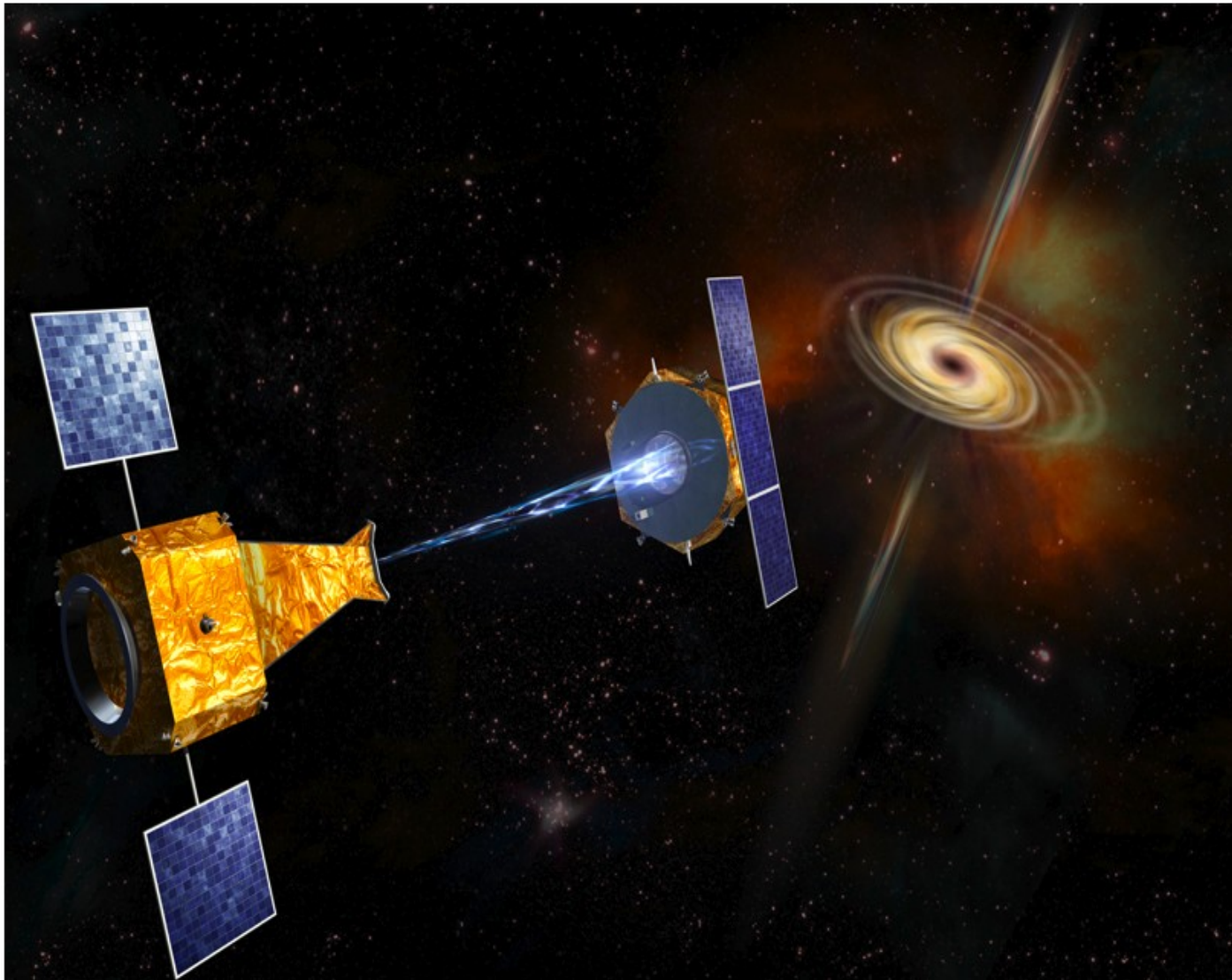


# Simbol-X: an Overview



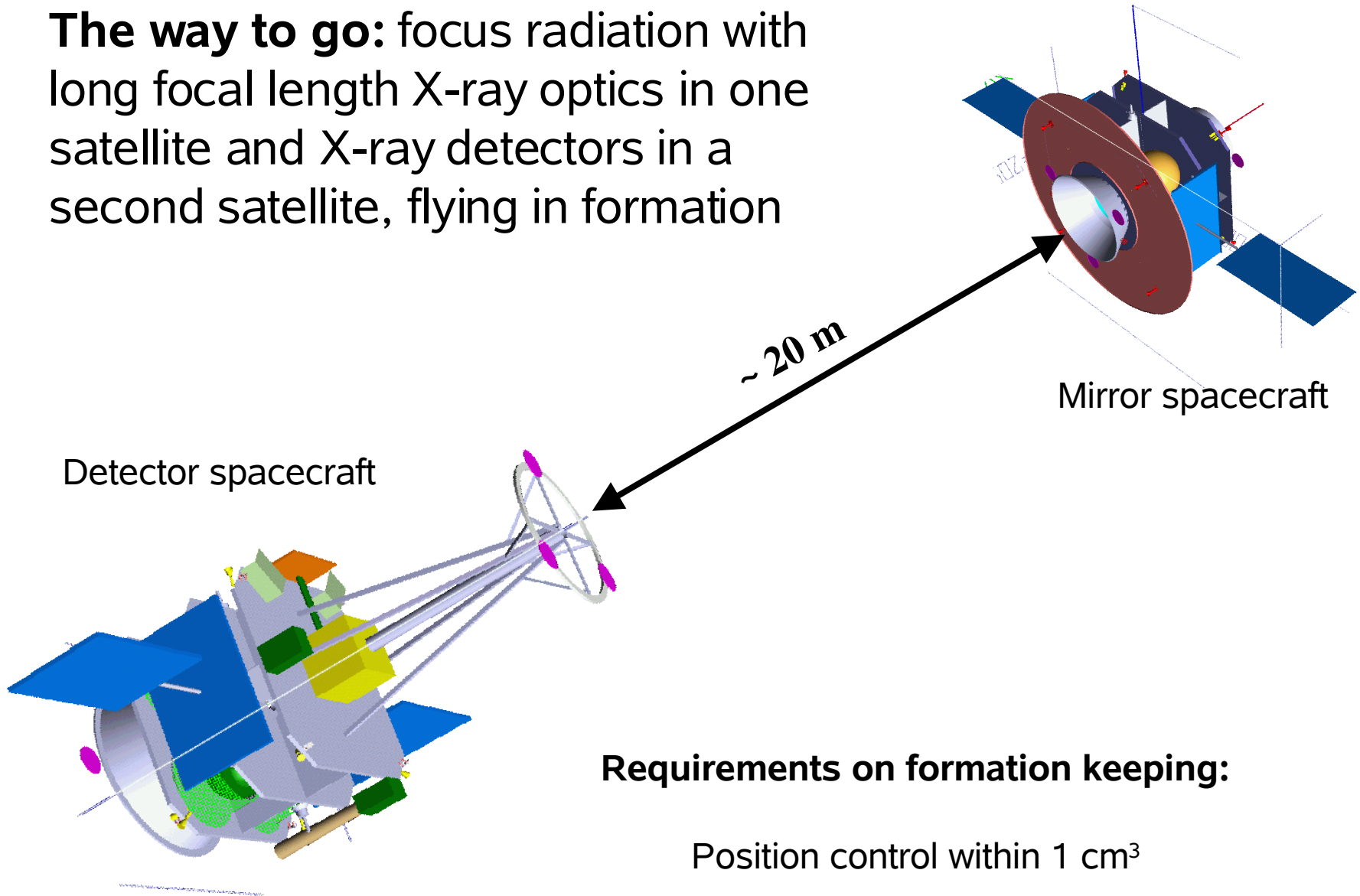
## History and current context of the project

- Since 2001, discussions between French and Italian X-ray astronomers about a new X-ray mission in the 0.1–80 keV range.
- Programmatic contacts have been initiated between CNES and ASI.
- ASI and CNES agree to perform a joint feasibility study (phase A) on a X-ray mission called Simbol–X.
  - MPE and IAATübingen were invited to contribute to Simbol–X  
**Contribution: Low Energy Detector (LED)**
    - Large French and Italian Consortium, including:  
CEA/Saclay, APC/Paris, CESR/Toulouse, and  
INAF-OAB/Milano, INAF/Bologna, INAF/Roma, INAF/Palermo
- **Simbol–X will offer an improvement of over two orders of magnitude in sensitivity and a much better angular resolution compared to the instruments currently operating above 10 keV.**

## Summary of Simbol-X top-level scientific requirements

Parameter	Requirement
<i>Energy band:</i>	0.5 – $\geq$ 80 keV
<i>Field of view (at 30 keV):</i>	$\geq$ 12' (diameter)
<i>On-axis effective area:</i>	$\geq$ 100 cm <sup>2</sup> at 0.5 keV $\geq$ 1000 cm <sup>2</sup> at 2 keV $\geq$ 600 cm <sup>2</sup> at 8 keV $\geq$ 300 cm <sup>2</sup> at 30 keV $\geq$ 100 cm <sup>2</sup> at 70 keV $\geq$ 50 cm <sup>2</sup> at 80 keV (goal)
<i>Detectors background</i>	$< 2 \times 10^{-4}$ cts s <sup>-1</sup> cm <sup>2</sup> keV <sup>-1</sup> HED $< 3 \times 10^{-4}$ cts s <sup>-1</sup> cm <sup>2</sup> keV <sup>-1</sup> LED
<i>On-axis sensitivity</i>	$\leq 10^{-14}$ c.g.s. (~0.5 $\mu$ Crab), 10-40 keV band, 3 $\sigma$ , 1Ms, power law spectrum with $\Gamma=1.6$
<i>Line sensitivity at 68 keV</i>	$< 3 \times 10^{-7}$ ph cm <sup>2</sup> s <sup>-1</sup> , 3 $\sigma$ , 1Ms ( $2 \times 10^{-7}$ goal)
<i>Angular resolution</i>	$\leq 20''$ (HPD), E < 30 keV $\leq 40''$ (HPD) @ E = 60 keV (goal)
<i>Spectral resolution</i>	E/ $\Delta$ E = 40-50 at 6-10 keV E/ $\Delta$ E = 50 at 68 keV (goal)
<i>Absolute timing accuracy</i>	100 $\mu$ s (50 $\mu$ s goal)
<i>Time resolution</i>	50 $\mu$ s
<i>Absolute pointing reconstruction</i>	$\sim 3''$ (radius, 90%) (2'' goal)
<i>Mission duration</i>	3 years including commissioning and calibrations (2 years of scientific program) + provision for a possible 2 year extension
<i>Total number of pointings</i>	$> 1000$ (first 3 years, nominal mission) 500 (during the possible 2 year mission extension)

**The way to go:** focus radiation with long focal length X-ray optics in one satellite and X-ray detectors in a second satellite, flying in formation

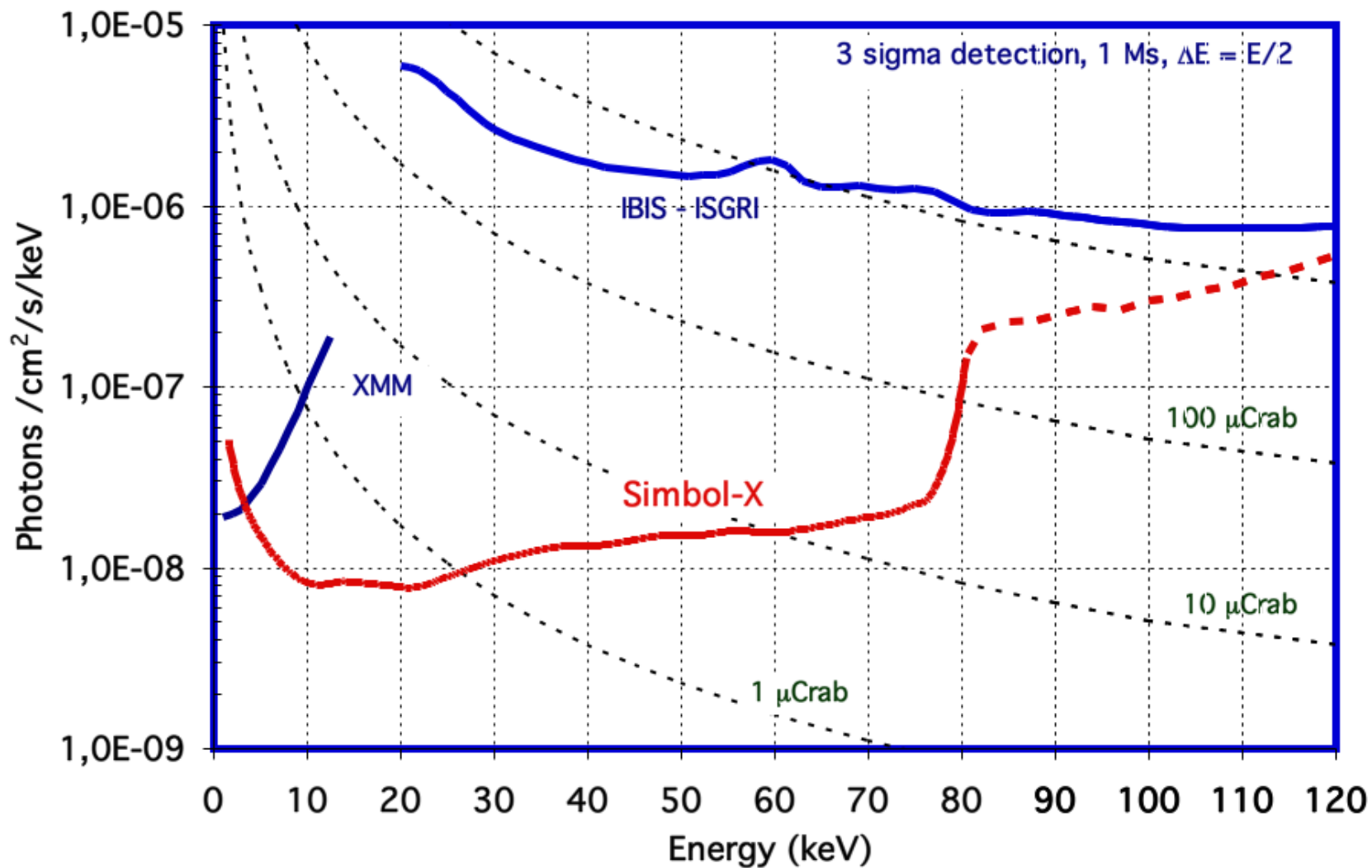


**Requirements on formation keeping:**

Position control within 1 cm<sup>3</sup>

# SIMBOL-X Sensitivity

Focal length 20 m, Field Of View 12 arcmin

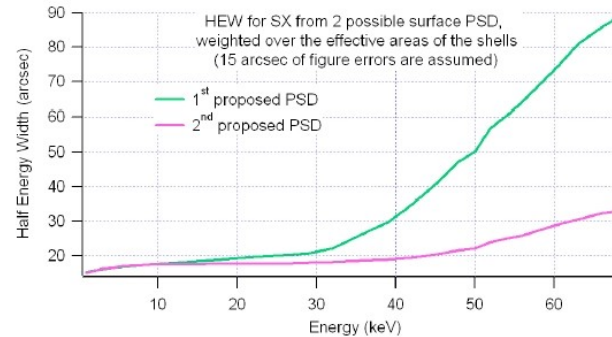
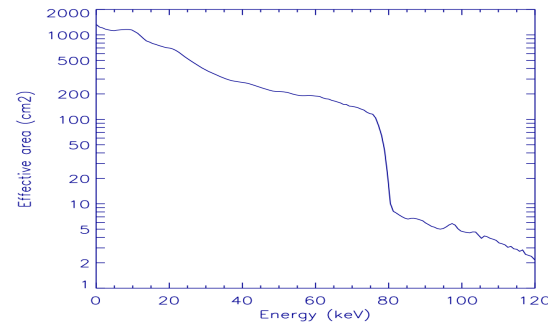
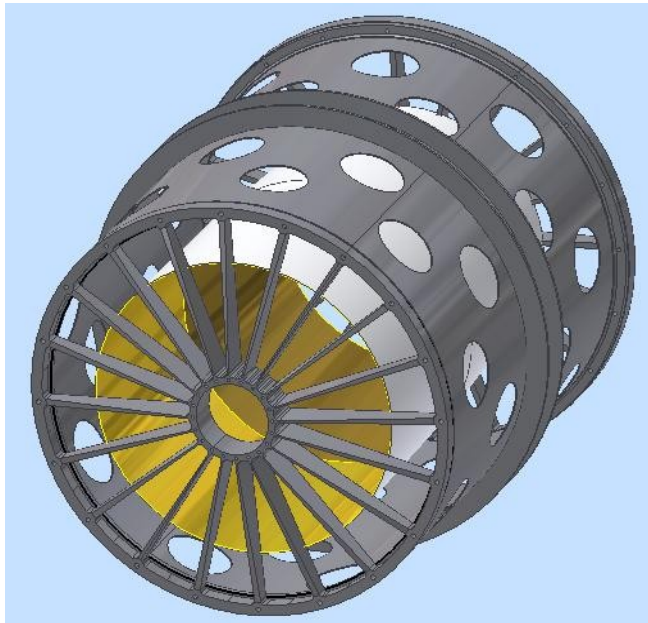


# Optics (OAB / Media Lario)

G. Pareschi et al.

## Parameters

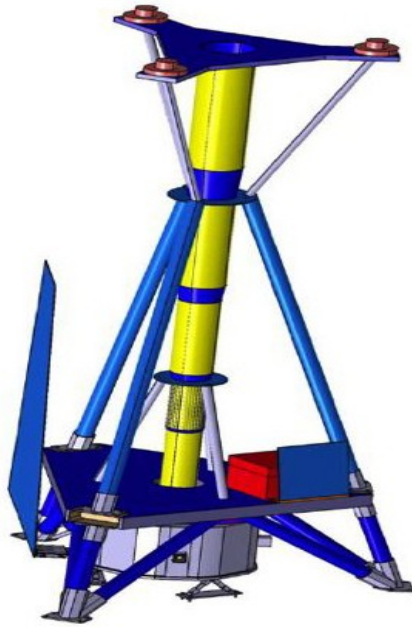
- 100 shells (Ni, replicated from mandrels, XMM)
- focal length : 20 m
- max diameter : 70 cm
- shell thickness ~ 2-3 times less than XMM
- Pt/C multilayer (sputtering inside mirror shells)



**Mirror Tests and Calibration in Neuried at the Panter Facility**  
**Modification of the Facility necessary**

## Detector payload

Designed to ensure low internal background and rejection of the cosmic diffuse background, with good spectro-imaging capabilities.

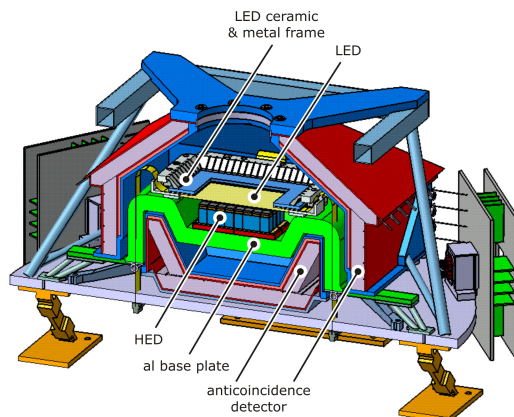


### Background shielding

- Graded shield collimator, to block diffuse background (together with sky-shield on mirror spacecraft)
- Active anticoincidence shield, surrounding imaging detectors

### Imaging detectors

- Two « sandwiched » systems, with a transition energy at 17 keV
- Both detectors are covering  $8 \times 8 \text{ cm}^2$  with  $625 \text{ }\mu\text{m}$  pixels
- Flexible read-out
- operating at moderately low temperature ( $\sim -40^\circ\text{C}$ )

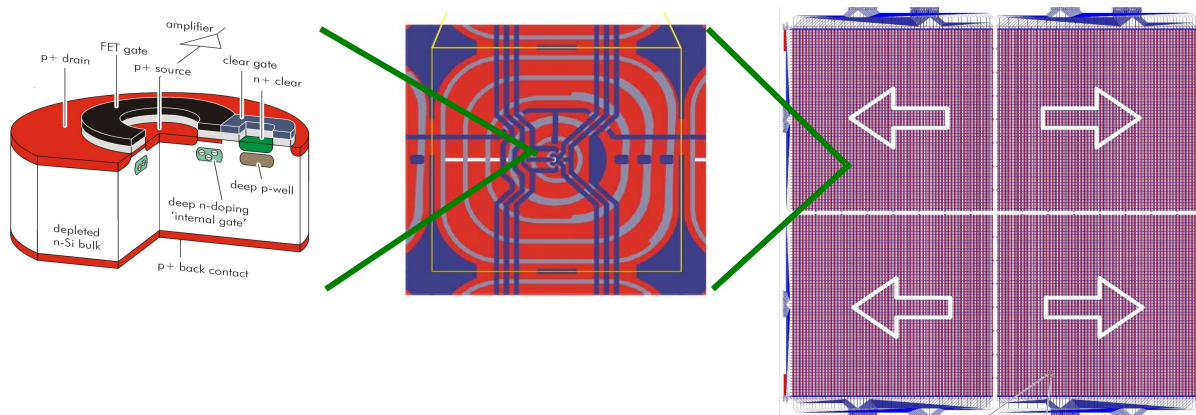


# Low Energy Detector (LED)

Peter Lechner, et. al

HLL(PNsensor)/IAAT/MPE/TUD/FAU

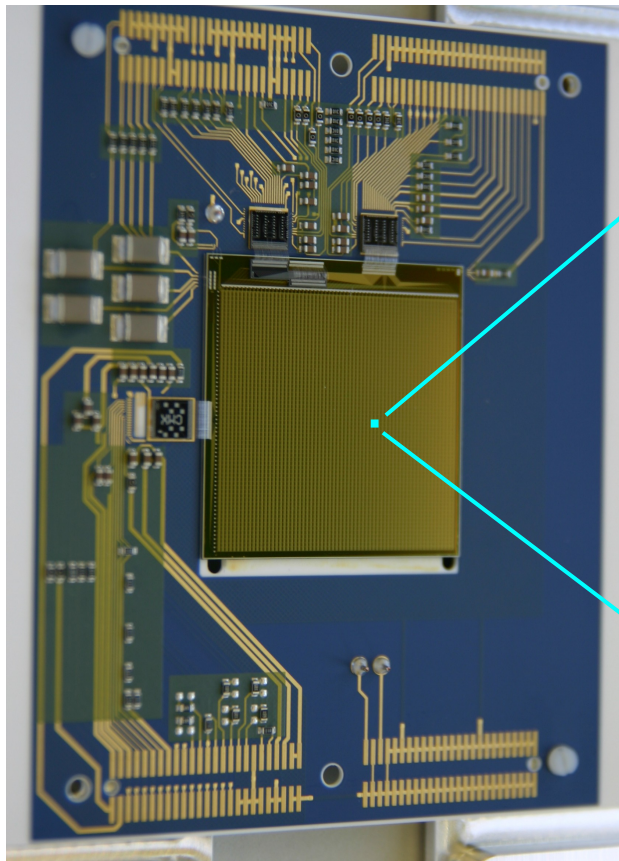
- Detector: Silicon Drift Detector with DEPFET readout  
**128 x 128** pixel, pixel size 625  $\mu\text{m}$   
divided into 4 independent quadrants  
readout time 128  $\mu\text{s}$ /frame in Fullframe mode  
(50  $\mu\text{s}$  in Window mode)  
spectral resolution: < 150 eV at 6 keV
- Event Pre-Processor: One unit per quadrant
- One single cold redundant Interface Controller



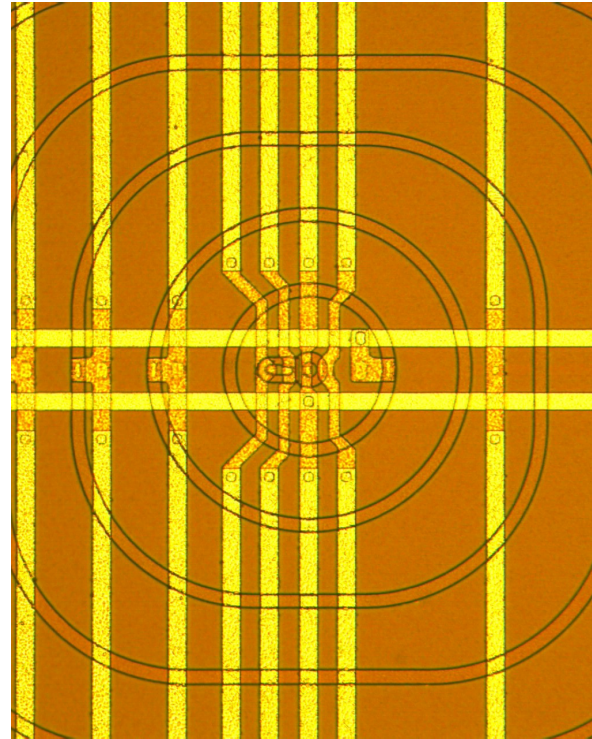


# LED quadrant test

■ SX-LED quadrant hybrid



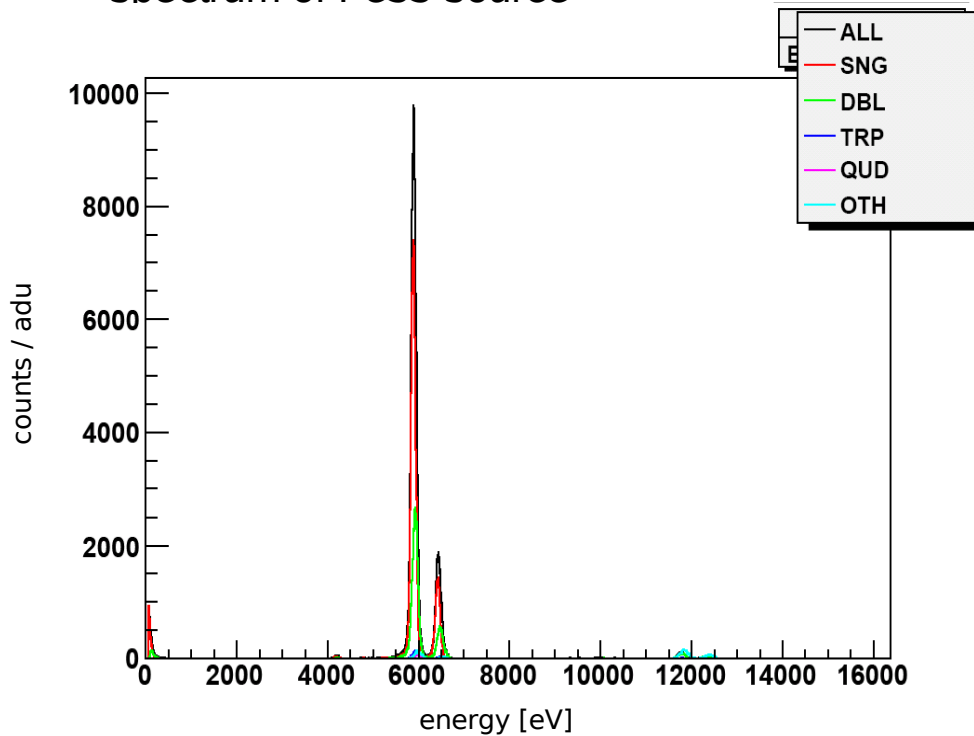
■ SX-LED pixel 500 x 500  $\mu\text{m}^2$



- sensitive area  $> 10 \text{ cm}^2$
- largest device since XMM-pnCCD

# LED quadrant test

■ spectrum of Fe55 source



■ voltages/currents

moderate values

- back -80 V
- drift ring -40 V
- drain 20  $\mu\text{A}/\text{pix}$

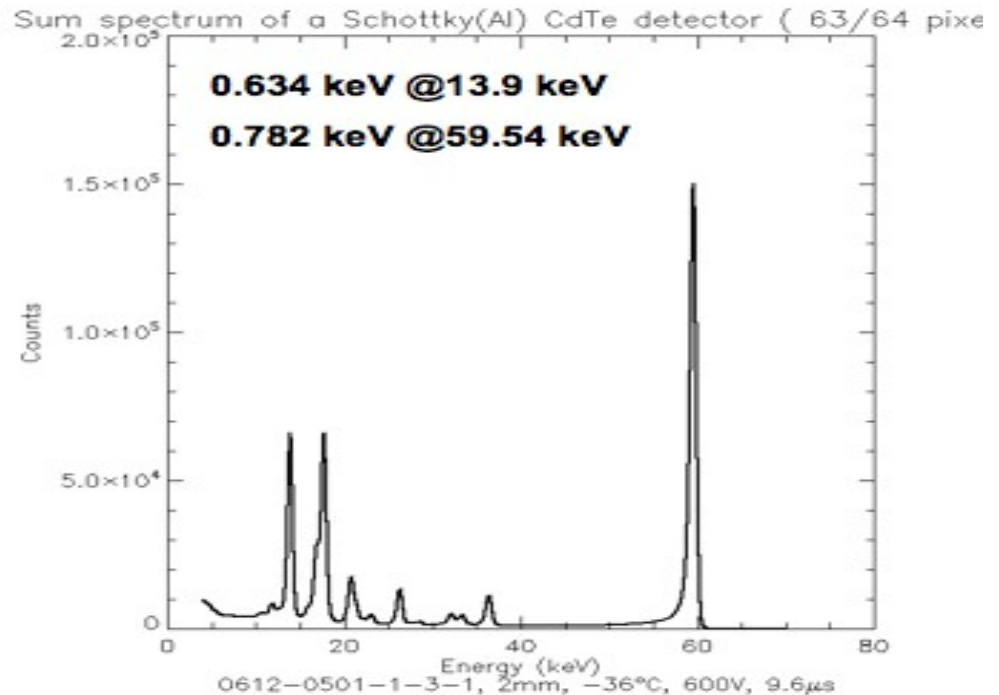
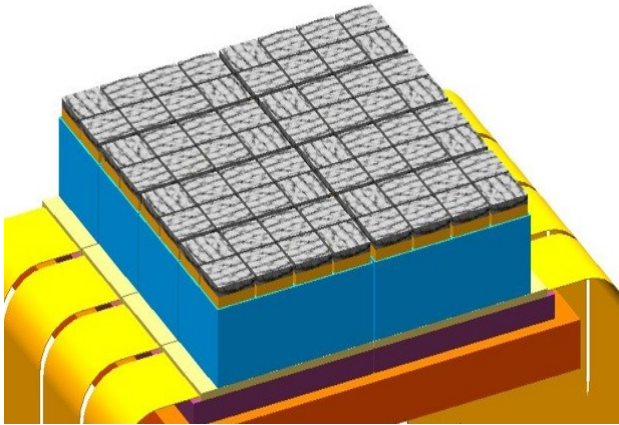
energy spectrum from flat-field illumination

- 127 eV FWHM @ 5.9 keV (single pixel events)
- peak/background ratio 3.000:1

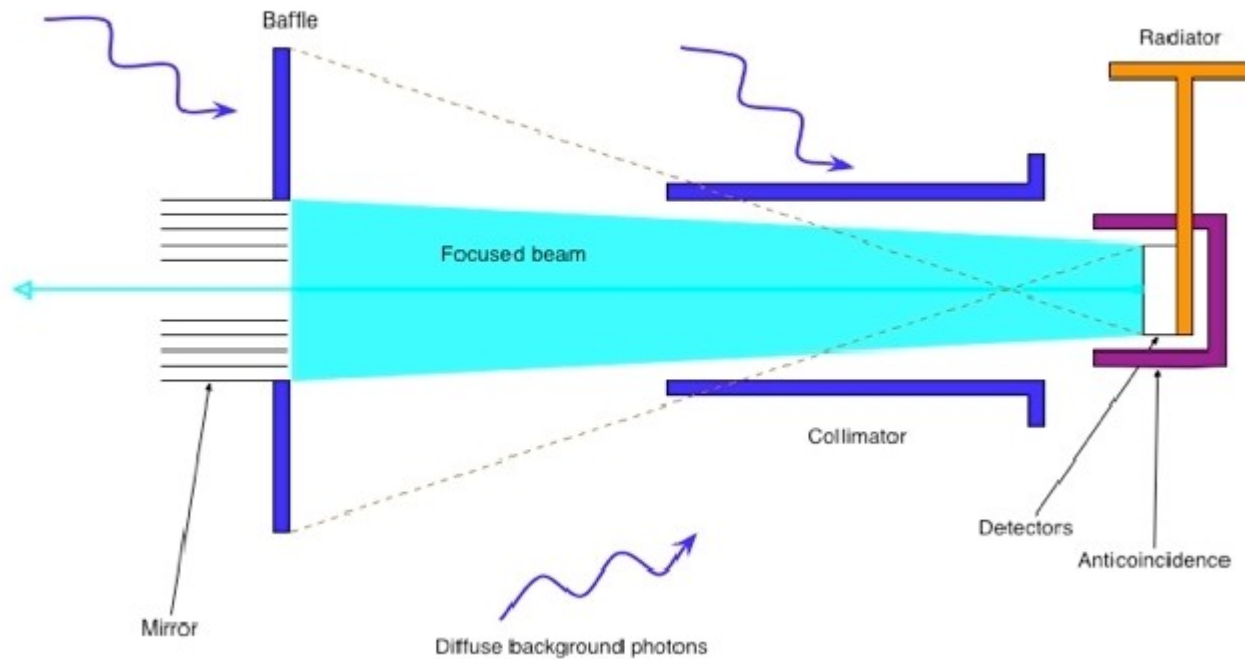
# High Energy Detector (HED)

(O. Limousin, P. Ferrando, et al., CEA Saclay)

- array of 64 pixellated Cd(Zn)Te modules of 256 pixels each, 2 mm thick,
- each module  $\sim 1 \text{ cm}^2$  total area, four side buttable
- with integrated front-end ASICs.
- self trigger read-out.
- spectral resolution:  $\sim 1 \text{ keV}$  at 60 keV



# Schematic view of the two satellites with baffle, collimator and anti-coincidence counter (AC)

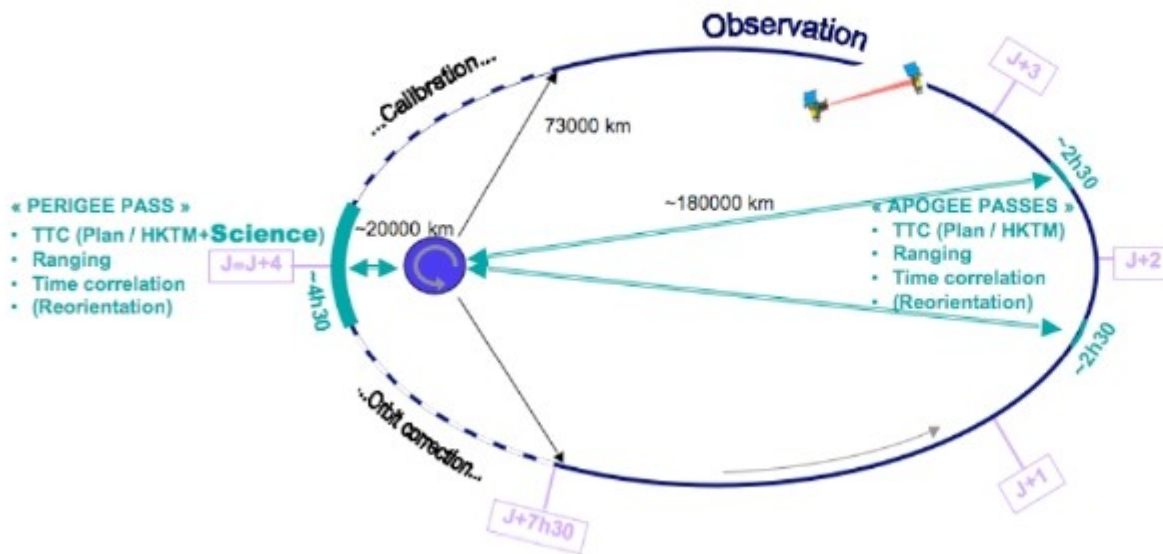


Detailed model of baffle, collimator, detectors and AC was basis for extensive Monte Carlo simulations of the background, using GEANT4

- done by R. Chipaux (CEA), **C. Tenzer (IAAT)** and M. Kuster (TUD)

# Mission Implementation

- Launch date: in the middle of 2014
- Mission parameters are :
  - single launch, by Soyuz from Kourou
  - high elliptical 4 day orbit (~ 20,000 – 180,000 km)
  - mission duration: 3 years with provision for a 2 years extension



Perigee :	20,000 km
Apogee :	180,000 km
Period :	4 days
Cycling ratio :	83 %
Time > 73,000 km :	~ 290 ks

**END**

